# An Experimental Study of Footprint Scale Variability of Raindrop Size Distribution





Ali Tokay (JCET-University of Maryland Baltimore County and NASA-Goddard Space Flight Center, Greenbelt, Maryland, USA) Leo Pio D'Adderio (Department of Physics and Earth Science, University of Ferrara, Ferrara, Italy) Federico Porcù (Department of Physics and Astronomy, University of Bologna, Bologna, Italy) David B. Wolff (NASA-GSFC Wallops Flgith Facility, Wallops Island, VA) and Walter A. Petersen (NASA Marshall Space Flight Center, Huntsville, AL

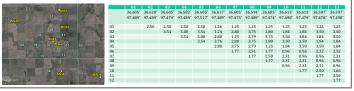


## 1. Introduction

Non-Unform Beam Filling (NUBF) within footprint of the dual frequency precipitation radar (DPR) on board NASA's Global Precipitation Measurement (GPM) mission core satellite is one of the key uncertainties of the precipitation retrieval algorithms. NUBF is a combined effect of precipitation gradient and intermittence within the footprint and occurs due to spatial variability in both horizontal and vertical direction. While the scanning and vertically pointing radars are the solo source to study the spatial variability in vertical, the dense rain gauge and disdrometer networks are often employed to determine the spatial variability in horizontal direction. The disdrometer networks are mainly available through field campaigns and in a few ground validation sites but are superior to the gauge network due to their higher temporal resolution and ability to measure raindrop size distribution and related integral rain parameters. This study uses disdrometer network that was operated during from Mid-latitude Continental Convective Clouds (MC3E) campaign to study the spatial variability within the ropornitor DPR.

# 2. Field Campaign

The MC3E campaign, a joint effort by the US Department of Energy's Atmospheric Radiation Measurement (ARM) and the NASA GPM GV programs, was conducted in North Central Oklahoma (36.7N, 97.1W) from April 22 to June 6, 2011. Seven (five GPMGV and two ARM) third-generation compact two-dimensional video disdrometers (2DVD) were deployed at and around the ARM Southern Great Plains site where the distances ranged from 0.4 to 9.2 km. For the purpose of this study, it would have been desirable to have uniformly distributed units within a circular area with a diameter of 5 km, representing the DPR footprint. This was not feasible due to power and open space requirements of the site selection. However, the layout and number of units allowed interpolating the 2DVD measurements tot desired points. Figure 1 shows the locations and interpolation points of the 2DVDs. The locations are given based on serial number (SN), while the interpolated points are numbered from 01 to 13. Table 1 presents the coordinates and distances between the points.



### 3. Data Analysis

The standard processing by GPMGV defines rainy minutes as having a minimum rain rate of 0.01 mm h<sup>-1</sup> and a minimum number of drops of 10 sampled in one-minute observations. This resulted in an MC3E database or 740 one-minute samples, which were the input for the interpolation routine. Rain intermittence is frequently observed in light stratiform rain and plays an important role in NUBF. We addressed this matter by setting DSI derived rain rates less than 0.1 mm h<sup>-1</sup> as zero after interpolation. This eliminated 5 samples where the rain rate fell below the threshold in all 13 sites. Surprisingly, only 34 samples had zero rainfall at one or more sites for the remaining dataset. Since the DPR footprint is a single data point from GPM perspective, the areal average rainfall is considered to determine the rainy footprint. Hence, a rain rate threshold of 0.1 mm h<sup>-1</sup> was applied to the areal average rainfall and 12 more samples were eliminated retaining 723 samples. At least 4 out of 13 site had zero rainfall in these 12 samples even though one sample had 5 zero rainfall sites but areal average rainfal was above the threshold. Post GPM launch, DPR minimum detectable signals are currently establised at 13 dBz for the Ka-band high sensitive swath (HS) and for Ku-band normal swath (NS) and 18 dBZ for the matched Ka and Ku-band swath (MS). The corresponding sample sizes were 703 and 698 for HS and NS, respectively, and were 639 for MS.

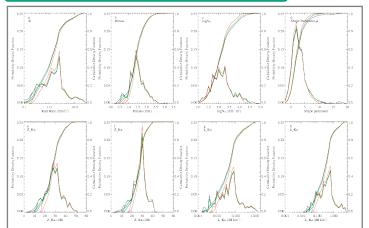
## 4. Methodology

A three-parameter exponential function is adopted to investigate the spatial variability of DSD and integral rai parameters. The exponential function is expressed as:

$$r(d) = r_0 \exp\left(-\frac{d}{d_0}\right)^{s_0}$$

where ro, so are nugget and shape parameters, respectively, and do is the correlation distance. The Pearson correlation coefficient, r, is calculated between the paired 2DVD observations at distance, d. The ro is the correlation between the collocated observations and is set to 0.99 in the absence of collocated 2DVDs. It should be noted that there is variability in drop counts at spatial scales 1-100 m but the correlations of derived DSD and integral rain parameters are mainly higher than 0.90. An initial guess was made for do and so using ranges of 0 to 300 at an increment of 0.1 and 0 to 2 at an increment of 0.01, respectively. The do and so are calculated minimizing the root-mean square error (RMSE) between the observation and equation based correlations. The RMSE is the measure of the goodness of the fit and it is critical for the interpretation of do and so.

# 5. Probability and Cumulative Distributions



Moderate-to-heavy rain in southern Great Plains receives relatively high percent of contribution from large drops (> 3 mm in diameter). The presence of large drops results in 14% of the observations  $D_{mass} \ge 2.0$  mm a MS threshold. For the same threshold, 11% and 16% of the observations had  $Z_Ku \ge 40$  dBZ and  $k_i Ka \ge 1$  dB km<sup>-1</sup>, respectively. The presence of large drop results also in broader size distribution with relatively low shape parameter of the modeled gamma distribution. The MC3E dataset exhibited drastically differen properties than a preceding study conducted at the mid-Atlantic coastal site of Wallops Island, Virginia where virtually no observations of R  $\ge$  10 mm h<sup>-1</sup>,  $D_{mass} \ge 2.0$  mm, and  $Z_i Ku \ge 40$  dBZ occurred.

### 6. Inverse Distance Weight

The accuracy of the Inverse Distance Weight (IDW) is tested through cross comparison of four physica parameters,  $D_{mass}$ ,  $D_{gNw}$ , R, and Z, Ku, between SN46 and site #01 and between SN47 and site #01. Site #01 is 0.15 and 0.34 km from SN46 and SN47, respectively. There were 707 samples where SN46 and SN47 measured  $R \ge 0.1$  mm  $h^{-1}$ . Figure 2 reveals an excellent agreement between the observed and interpolated parameters. There was also no systematic over- or under-estimation of any parameter. Site #01 had a bette agreement with SN46 due to its closer distance. Rain rate was slightly overestimated with 0.6% and 1.8% bias and 4% and 20% absolute bias with respect to SN46 and SN47, respectively.

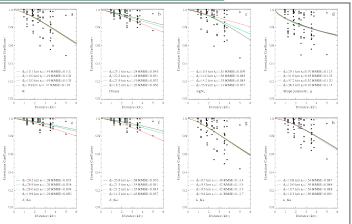






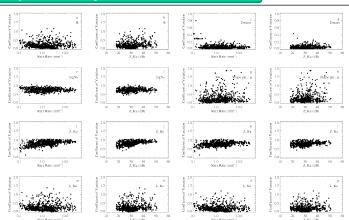


### 7. Spatial Variability: Correlation Coefficient



The correlations at a given distance were not sensitive to the choice of threshold for R, k\_Ku and k\_Ka fields. The parameters of exponential function were therefore very close to each other for all four thresholds. At the same time, there were substantial differences in correlations at a given distance. The correlation of R ranged from 0.93 to 0.48 at 5 km distance for a given R threshold. This resulted in high RMSE for these three fields. The correlations of  $\mu$  were not sensitive to the choice of threshold but highly sensitive to the directional variability. The RMSE of the fit was therefore the greatest for  $\mu$  among all fields. The so was lower than 1.0 only for this field and corresponding do was high even though the correlations were less than 0.8 for nearly half of the observations. The correlations at a given distance were more sensitive to the choice of threshold for Z, Kz, Ka, and D\_mls-fields where correlations remained above 0.75 regardless of distance in R threshold database. The correlations of logNw also showed sensitivity to the choice of threshold but in reverse order. The exponential fits were the lowest and highest RMSE at MS and R thresholds, respectively.

### 8. Spatial Variability: Coefficient of Variation



Coefficient of Variation (CV) represents the degree of uniformity of the parameter from highly uniform (CV  $\lesssim 0.25$ ) to extremely variable (CV > 1.00) classes. The remaining three classes may be called mostly uniform (0.25 < CV < 0.5), mostly (or moderately) variable (0.5 < CV < 0.75), and highly variable (0.75 < CV < 1.0). All physical parameters had CV < 1.25 except the shape parameter of gamma model distribution. The vast majority of R had CV < 0.5, only 8% of the observations showed moderate to extreme variability (CV > 0.5). Dmsss was highly uniform across the range of R and Z\_Ku values and had two outliers with CV of 0.7 and 0.8. These outliers occurred in back-to-back minutes during a deep convective event where one and two sites did not report rainfall. The vast majority of the observations fell into moderate to high variability classes for CV of logNw. The shape parameter, on the other hand, had over a dozen samples where CV > 2 was (not shown). These samples correspond to DSD where gamma distribution is probably not the best choice. The CV of Z\_Ku exhibited mostly high variability but significant number of samples had also moderate variability. The CV of k\_Ku fell mostly into the highly uniform class and the majority of the database had CV < 0.5.

### Acknowledgments

Thanks to Patrick Gatlin and Matt Wingo for maintenance of 2DVD during MC3E. Acknowledgments extend to V.N. Bringi and Merhala Thurai for their efforts in calibrating 2DVD at the beginning of the field campaign. We appreciate DOE ARM for hosting the field campaign in their Southern Great Plains facility and providing two 2DVD database. This research was funded through NASA grant NNX16AD45G. Funding support from the GPM Grould Validation Program and NASA PMM Science Team funding provided by Dr. Ramesh Kakar are greatly